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# **High-frequency Earth rotation variations deduced from** altimetry-based ocean tides

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Abstract A model of diurnal and semi-diurnal variations in Earth rotation parameters (ERP) is constructed based on altimetry-measured tidal heights from a multi-mission empirical ocean tide solution. Barotropic currents contributing to relative angular momentum changes are estimated for nine major tides in a global inversion algorithm that solves the two-dimensional momentum equations on a regular 0.5° grid with a heavily weighted continuity constraint. The influence of 19 minor tides is accounted for by linear admittance interpolation of ocean tidal angular momentum, although the assumption of smooth admittance variations with frequency appears to be a doubtful concept for semi-diurnal mass terms in particular. A validation of the newly derived model based on post-fit corrections to polar motion and universal time  $(\Delta UT1)$  from the analysis of Very Long Baseline Interferometry (VLBI) observations shows a variance reduction for semi-diurnal  $\Delta$ UT1 residuals that is significant at the 0.05 level with respect to the conventional ERP model. Improvements are also evident for the explicitly modeled K<sub>1</sub>, Q<sub>1</sub>, and K<sub>2</sub> tides in individual ERP components, but large residuals of more than 15 µas remain at the principal lunar frequencies of O1 and M2. We attribute these shortcomings to uncertainties in the inverted relative angular momentum changes and, to a minor extent, to violation of mass conservation in the empirical ocean tide solution. Further dedicated hydrodynamic modeling efforts of these anomalous constituents

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are required to meet the accuracy standards of modern space geodesy.

**Keywords** Earth rotation variations · Empirical ocean tides · Tidal currents · Angular momentum changes · VLBI

# **1** Introduction

A large fraction of high-frequency Earth rotation variations is due to diurnal and semi-diurnal ocean tides. Water masses are redistributed in response to the tidal forces exerted by the Moon and the Sun, generating relative angular momentum with respect to the rotating reference frame and causing sea surface changes that influence Earth's inertia tensor. These two mechanisms are classically referred to as the motion and mass terms of ocean tidal angular momentum (OTAM, e.g., Chao and Ray 1997) and manifest themselves in perturbations of the Earth's rotation vector. Axial perturbations will affect the rotation rate and may be quantified either as changes in length-of-day ( $\Delta$ LOD) or increments to universal time ( $\Delta$ UT1), while equatorial (horizontal) displacements of the rotation axis from the body axis will give rise to polar motion  $\{x_n, y_n\}$  and nutation. The subset containing both polar motion and  $\Delta UT1$  (or  $\Delta LOD$ ) is commonly referred to as Earth rotation parameters (ERP).

For individual tidal constituents, ocean-induced ERP variations are at the level of a few hundred  $\mu$  as (microarcseconds); cf. Chapter 7 of Petit and Luzum (2010) and note that  $\Delta$ UT1 is usually reckoned in  $\mu$ s. Early evidence of these signals in geodetic Very Long Baseline Interferometry (VLBI) data by Brosche et al. (1991) and Herring and Dong (1991) was substantiated through refined VLBI and Satellite Laser Ranging (SLR) analyses in the mid-1990s (Sovers et al. 1993; Herring and Dong 1994; Watkins and Eanes 1994; Gipson

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1996) and prompted efforts of modeling the tidal effects in an a priori fashion, using hydrodynamic equations and knowledge of the tide-generating potential (Brosche et al. 1989; Seiler 1991; Gross 1993). Exploiting the global record of sea level observations from satellite altimetry, Ray et al. (1994) and Chao et al. (1996) finally documented a satisfactory agreement between space geodetic observations and geophysical model quantities.

At present, the International Earth Rotation and Reference Systems Service (IERS, Petit and Luzum 2010) recommends accounting for the high-frequency ocean tidal ERP variations through routines that build on the "Model C" results of Chao et al. (1996) as derived from Egbert et al. (1994) and Arctic OTAM supplements from Ray et al. (1997)<sup>1</sup>. Several authors have, however, found shortcomings in this model and emphasize the need for an updated version (Steigenberger et al. 2006; Artz et al. 2010; Böhm 2012; Griffiths and Ray 2013). To a minor degree, differences between the observed tidal amplitudes and the values prescribed by the IERS arise from the omission of radiational tidal effects in the coupled atmosphere-ocean system at the order of 5 µas (Brzeziński et al. 2002; Schindelegger 2014) and from slight imperfections (<3 µas Brzeziński 2003) in accounting for the semi-diurnal spin and diurnal polar motion librations of the Earth (Chao et al. 1991). Larger residuals in the range of 15–30 µas are evident for the major gravitational tidal lines (Böhm 2012) and suggest that the analysis of Egbert et al. (1994) was affected by uncertainties in the observed sea level elevations, a comparatively short altimetric dataset, errors in bathymetry charts, or simplified hydrodynamic equations.

One option for revising the current IERS standard is to construct an observation-based ERP model from the tidal variations sensed by VLBI, Global Navigation Satellite Systems (GNSS), or SLR. Several models of this type have been derived in the past and show unprecedented precision levels (Gipson 1996; Steigenberger et al. 2006; Artz et al. 2011). However, such datasets cannot serve as independent standards in the analysis of space geodetic observations and are usually marred by technique-specific systematic errors that leak into the tidal estimates (Griffiths and Ray 2013). In fact, models from different single techniques deviate significantly from each other (Artz et al. 2011). Artz et al. (2012) derived a combined VLBI/Global Positioning System (GPS) model which partly reduces these inconsistencies, but observations from GPS dominate the solution.

A more independent approach is to draw on the increasingly accurate account of ocean tides made possible by altimetry and hydrodynamic modeling; see Stammer et al. (2014) for a recent, comprehensive overview. Three categories of tide models exist and assume different roles in the context of Earth rotation. First, purely hydrodynamic (or forward) tide models are mere solvers of the shallow water equations, using only knowledge of bathymetry and the astronomical forcing. Ideas about drag and dissipation can be tested freely in such environments (Arbic et al. 2004), butwith the exception of the atmosphere driven  $S_1$  tide (Ray and Egbert 2004)-forward-modeled elevations and currents are generally too inaccurate to meet the requirements of space geodesy. Second, assimilation models combine hydrodynamic cores with sea level observations from altimetry in one common adjustment and feature root-mean-square (RMS) misfits of less than 0.6 cm per constituent when validated against ground truth estimates from deep-water bottom pressure stations (Stammer et al. 2014). Accordingly, recent data-assimilative models have been shown to perform well in the analysis of VLBI observations (Böhm 2012). Future developments in that field are thus likely to aid the description of high-frequency signals in Earth rotation. Last, tide models can be also empirical, derived, e.g., from a least-squares harmonic analysis of altimetric measurements without invoking hydrodynamic equations that entail inevitable simplifications. While these models possibly provide the most accurate realization of the true tide in terms of elevation, they do not include any information about the associated deep-water currents which are required for the evaluation of the motion term contribution to OTAM. Yet, using a full elevation grid as constraints, barotropic (2D) currents can be inferred from an inversion of the linearized shallow water equations (Ray 2001).

The main goal of this paper is to deduce and validate an independent model for diurnal and semi-diurnal ERP variations based on EOT11a (EOT for short, Savcenko and Bosch 2012), a state-of-the-art empirical ocean tide model derived at Deutsches Geodätisches Forschungsinstitut from multimission satellite altimetry data. Mission-specific weightings are based on a variance component estimate and the residual harmonic analysis up to latitudes of  $\pm 81.5^{\circ}$  was performed relative to FES2004 (Lyard et al. 2006). In a comparison to ground truth estimates from 151 deep-ocean gauges (Stammer et al. 2014), EOT has been shown to yield reasonably good RMS statistics, similar to those of other contemporary empirical and data-assimilative tide models. A systematically better performance of EOT that would also endorse our model choice in the present work is, it might be admitted, not apparent. We nonetheless emphasize that the network of instruments for the tide determinations in Stammer et al. (2014) is unevenly distributed across the globe and cannot serve as a measure of the broad-scale elevation features that couple to Earth rotation signals (degree-two spherical harmonics). We thus expect EOT to be a reliable source to advance our knowledge of high-frequency ocean tidal ERP variations.

<sup>&</sup>lt;sup>1</sup> The work of Ray et al. (1994) is often erroneously cited as the source model of the IERS tables; see http://bowie.gsfc.nasa.gov/ggfc/tides/ (accessed 20 March 2016) for a dementi

Having said that, it is also important to set out quantitative measures by which the success of our modeling efforts can be evaluated. Griffiths and Ray (2013) provide evidence that the spurious harmonic aliases seen in nearly all GNSS products are reasonably well simulated if one assumes error amplitudes of 20% in the conventional ocean tidal ERP model. This estimate of uncertainty is somewhat larger but generally consistent with the spread of ERP tidal corrections from different ocean tide datasets in Böhm (2012). Reducing the present error budget of the IERS model to a level of only a few percent—and thus guaranteeing an accuracy of 1 mm for the deliverables of space geodesy-should serve as an ultimate (though ambitious) goal for studies as the one presented here. Even if these requirements are not adequately met, the method of using and inverting a modern multimission altimetric dataset for the derivation of tidal ERP corrections can point the way forward toward achieving the desired geodetic goals.

We have organized the paper in the following manner. The next section describes the algorithm by which ocean tidal currents are computed from measured elevations, before the treatment of minor tidal effects is discussed in Sect. 3. We validate our ERP model in an analysis of 20-year VLBI observations (Sect. 4) and summarize our results in Sect. 5.

# 2 Derivation of ocean tidal currents

To derive the flow field from empirical tidal heights, this work uses an inversion method described by Ray (2001). Instead of full hydrodynamic approaches that assimilate altimetry observations into the model, the measured elevations are kept fixed in the estimation procedure, and the currents are inferred from a gradient technique supplemented by a continuity constraint. This approach overcomes several shortcomings of the simple pointwise gradient estimation (Ray 2001), but also transfers the local determination of currents into a global inversion process.

## 2.1 Dynamic equations

We essentially assume shallow water dynamics with advection and turbulent viscosity terms omitted. The two-dimensional momentum equations in this case are (Ray 2001)

$$\boldsymbol{u}_{t} + f\boldsymbol{k} \times \boldsymbol{u} + \frac{\boldsymbol{F}}{\rho_{w}D} = -g\nabla(\zeta - \zeta_{\mathrm{E}} - \zeta_{\mathrm{SAL}}), \qquad (1)$$

and the continuity equation reads

$$\zeta_t = -\nabla(D\boldsymbol{u}),\tag{2}$$

with  $u_t$  being the time derivative of the horizontal tidal flow u = (u, v) and  $f = 2\Omega \sin \varphi$  being the Coriolis parameter,

where  $\Omega$  is the mean angular velocity of the Earth and  $\varphi$  is the geographic latitude.  $\mathbf{k}$  is a unit vector pointing to the local vertical,  $\rho_w$  is the nominal seawater density, D represents the water depth, and g is the gravitational acceleration.  $\nabla$  (nabla) is the spherical del operator, denoting horizontal derivatives. The tidal height  $\zeta$  is reckoned relative to the moving seabed and has to be corrected for equilibrium tidal forcing ( $\zeta_E$ ) and self-attraction and loading ( $\zeta_{SAL}$ ), both of which are described below. The stress term  $\mathbf{F}$  in our linearized model consists of a simple bottom friction of the form (Egbert et al. 1994)

$$F_{\lambda} = \kappa \rho_{w} u$$

$$F_{\varphi} = \kappa \rho_{w} v.$$
(3)

Dissipation schemes accounting for internal wave generation over rough topography (Green and Nycander 2013 2.b and 2.c, *ibid.*) were tested, but resulted in angular momentum deviations of about 10–20% with respect to established OTAM reference values. In contrast, a uniform frictional parameter led to more satisfactory results that were only weakly dependent on the exact value of  $\kappa$ . Following Ray (2001), we finally adopted  $\kappa = 0.01 \text{ ms}^{-1}$  for all of our computations.

Amplitudes (*A*) and phases ( $\phi$ ) of each tidal constituent were converted to complex amplitudes  $\hat{\zeta} = A_{\zeta} e^{-i\phi_{\zeta}}$ ,  $i \equiv \sqrt{-1}$  and expressed as  $\zeta = \hat{\zeta} e^{i\omega t}$ , where  $\omega$  is the tidal frequency. Thus, Eqs. (1) and (2) can be written as (Ray 2001; Logutov and Lermusiaux 2008)

$$i\omega\hat{U} - f\hat{V} + \frac{\kappa\hat{U}}{D} = -\frac{gD}{R_E\cos\varphi}\frac{\partial}{\partial\lambda}(\hat{\zeta} - \hat{\zeta}_{\rm E} - \hat{\zeta}_{\rm SAL}) \quad (4)$$

$$i\omega\hat{V} + f\hat{U} + \frac{\kappa\hat{V}}{D} = -\frac{gD}{R_E}\frac{\partial}{\partial\varphi}(\hat{\zeta} - \hat{\zeta}_{\rm E} - \hat{\zeta}_{\rm SAL})$$
(5)

$$\frac{1}{R_E \cos\varphi} \left[ \frac{\partial \hat{U}}{\partial \lambda} + \frac{\partial (\hat{V} \cos\varphi)}{\partial\varphi} \right] = -i\omega\hat{\zeta}, \tag{6}$$

where  $\hat{U} = D\hat{u}$  and  $\hat{V} = D\hat{v}$  are barotropic volume transports in the east and north direction, respectively,  $R_E$  is the mean radius of the Earth, and  $(\varphi, \lambda)$  denote latitude and longitude on a sphere. The bathymetry D in our numerical implementation below is a resampled 0.5° version of the 1-min General Bathymetric Chart of the Oceans (GEBCO) data, using moderate Gaussian weighting ( $\sigma = 0.5$ ).

## 2.2 Equilibrium tide and effects of SAL

Equilibrium tides represent the static response of a hypothetical, frictionless global ocean to differential gravitational forces. Neglecting third- and higher-degree terms,  $\zeta_E$  can be inferred from (Pugh and Woodworth 2014)

$$\zeta_{\rm E} = \gamma_2 \frac{V_{\rm tid}}{g},\tag{7}$$

where  $V_{\text{tid}}$  is the astronomical tide-generating potential (TGP) and  $\gamma_2 = 1 + k_2 - h_2$  is a frequency-dependent modifier to account for the body-tide deformation of the solid Earth using degree-two Love numbers  $k_2$  and  $h_2$  (Hendershott 1972). In Eq. (1),  $\zeta_E$  may be thus conceived as a proxy of the gravitational forcing of each tidal constituent, computed from tabulated equilibrium amplitudes  $A_E$  for either diurnal (subscript/argument d) or semi-diurnal (subscript/argument s) tides; cf. Table 1 of Arbic et al. (2004):

$$\zeta_{\mathrm{E},d}(\varphi,\lambda,t) = A_{\mathrm{E}}(d)\gamma_{2}(d)\sin(2\varphi)\cos(\omega(d)t+\lambda)$$
  

$$\zeta_{\mathrm{E},s}(\varphi,\lambda,t) = A_{\mathrm{E}}(s)\gamma_{2}(s)\cos^{2}\varphi\cos(\omega(s)t+2\lambda).$$
(8)

Partial tides not included in Arbic et al. (2004) were taken proportional to the respective potential from the HW95 catalog (Hartmann and Wenzel 1995). An initial (t = 0) phase lag for each equilibrium tide was introduced using the fundamental arguments of Simon et al. (1994).

The SAL tide represents a combination of feedback effects on the tidal dynamics that change sea level elevations at the order of 10% (Accad and Pekeris 1978). On the one hand, the redistributed water masses lead to a yielding of the ocean bottom and the associated body deformations in turn alter the Earth's gravitational field. On the other hand, water is attracted toward the tidal maxima (Gordeev et al. 1977). The rigorous treatment of these effects requires a convolution of the tidal elevation with a proper Green's function (Hendershott 1972), or a computation using spherical harmonics (Pugh and Woodworth 2014). A rough solution can be obtained by introducing a simple scaling factor  $\beta$ , i.e.,  $\zeta_{\text{SAL}} = \beta \zeta$ , with  $\beta = 0.085$  usually cited as the optimal choice for the M<sub>2</sub> tide (Accad and Pekeris 1978). Such global values can, however, produce considerable errors, especially in shallow seas (Ray 1998). We therefore calculated the SAL tide from the associated convolution (Ray 1998)

$$\zeta_{\text{SAL}}(\varphi,\lambda) = \rho_w R_E^2 \iint \zeta(\varphi',\lambda') \mathcal{G}(\psi) \cos \varphi' \, \mathrm{d}\varphi' \, \mathrm{d}\lambda', \quad (9)$$

where  $\mathcal{G}(\psi)$  are Green's function values computed from a combination of Load Love numbers  $\{k'_n, h'_n\}$  (Farrell 1973) and Legendre polynomials depending on the angular distance  $\psi$  from the load. We used the tabulated values of  $\mathcal{G}(\psi)$  from Stepanov and Hughes (2004), who calculated Green's functions for a bell-distributed mass instead of point masses to overcome the singularity at  $\psi = 0$ .

## 2.3 Interlude: OTAM computation

Much of our argumentation below involves OTAM quantities, computed as 2D integrals of tidal heights and volume transports from standard formulae given, e.g., in Chao and Ray (1997) (Eqs. 1 and 2, *ibid*.). Evaluation of the zonal mass term related to the inertia tensor increment  $\Delta I_{33}$  warrants special consideration, as it is intertwined with the question of mass conservation in global ocean tide models. Following Ray et al. (1997),  $\Delta I_{33}$  may be broken down into two orthogonal components

$$\Delta I_{33} = -\frac{2}{3}\rho_w R_E^2 \iint_O P_2(\sin\varphi) \zeta dO + \frac{2}{3}\rho_w R_E^2 \iint_O \zeta dO,$$
(10)

where  $P_2$  is the degree-two Legendre polynomial and Orepresents the ocean domain with surface element dO = $R_F^2 \cos \varphi d\varphi d\lambda$ . Tides in the real ocean do conserve mass; hence the total water mass change measured by the second term in Eq. (10) must be identically zero at every instant. In empirical or data-assimilative inverse models of the tides, though, strict mass conversation may be violated when fitting to observational data is prioritized over hydrodynamic information; see Egbert et al. (1994) on how to handle this trade-off in an inverse model. For EOT-as for some other models used in our study-the total water mass change of individual tides is indeed nonzero and perturbs estimates of  $\Delta I_{33}$  to an uncomfortably large extent, by up to 20% for the particular case of O<sub>1</sub>. This perturbation scales to only a few percent if  $\Delta UT1$  changes are evaluated from both inertia effects and the more dominant motion term contribution, but there is no guarantee that a dynamical inconsistency in the global elevation field leaves the associated (and numerically modeled) currents unaffected (Dickman 2010). While corrections of the tidal heights for nonconservation of mass lack clear physical guidelines, a bias in the axial inertia estimate may be avoided if only the first term in Eq. (10) is used to determine  $\Delta I_{33}$ ; see Ray et al. (1997) for further details. This procedure is followed throughout our study and does indeed reduce the spread of zonal mass term estimates from various ocean tide models.

## 2.4 Inversion algorithm

Equations (4)–(6) together with no-flow constraints at closed boundaries can be written as overdetermined linear system Ax = b, where A is a rectangular coefficient matrix, b is the right-hand side vector of observations, and x contains the unknown volume transports derived from a least-squares adjustment. The computations were performed on a staggered grid (C-grid, after Arakawa and Lamb 1977) at a resolution of 0.5° for both input and output data. Hence, the coefficient matrix consists of approximately 500,000 rows and 300,000 columns. The large number of zeros as well as the actual size of A make the use of iterative solvers feasible. Similar to Ray (2001) and Zahel (1995), we adopted the leastsquares algorithm *LSQR* for sparse matrices as published by Paige and Saunders (1982) and provided in MATLAB.

The convergence criterion for the least-squares solver is to compare  $\|\boldsymbol{b} - \boldsymbol{A}\boldsymbol{x}\| / \|\boldsymbol{b}\|$ , i.e., the norm of the relative residual in subsequent iterations. The algorithm stops (converges) when it falls below a user-defined tolerance parameter l. To find a reasonable value for l, a calibration experiment was designed based on HAMTIDE11a (HAM for short, Taguchi et al. 2014), an assimilation model for which both tidal elevations and currents are available. Inversions using several tolerances were calculated and converted to OTAM motion terms for comparison to angular momentum estimates from the original HAM flow field. Residuals were found to decrease down to  $l = 10^{-9}$  with very little improvement compared to  $l = 10^{-8}$ . However, for diurnal tides and  $l < 10^{-5}$ , residuals displayed a somewhat contrasting behavior across the three OTAM components, i.e., an improvement in the z-direction at the expense of increasing OTAM deviations in the y-direction. Hence, a threshold of  $l = 10^{-9}$ , as used in the calculations presented below, does not guarantee convergence for all constituents. In these cases, the solver was stopped after 200,000 iterations.

A diagonal matrix was used for preconditioning the coefficient matrix, making the  $L^2$  norm of all columns equal to one. With this modification, the iterative algorithm converges roughly in half the time. The results were also found to be slightly more accurate when the no-flow boundary conditions were down-weighted by a factor of 0.1. Finally, as discussed by Ray (2001), the continuity equation requires considerably larger weights than the momentum equations to obtain solutions of sufficient accuracy and to overcome singularities at critical latitudes around  $\pm 30^{\circ}$  for diurnal tides and  $\pm 75^{\circ}$  for semi-diurnal tides. Ray (2001) suggests a relative weight of c = 1000 for the continuity constraint, which is used as a benchmark value in the following section.

#### 2.5 Comparison with hydrodynamic solutions

A graphical validation of our inversion scheme is presented in Figs. 1 and 2, where volume transports deduced from HAM elevations of the  $K_1$  and  $M_2$  tides are compared to model-intrinsic (U, V) estimates. Global and regional flow patterns agree well among both solutions, but larger differ-



Fig. 1 Results of our tidal current inversion algorithm for  $K_1$ . The *right* column displays the estimated volume transports (eastward component U), while the *left column* shows the same component from the hydrody-

namic assimilation model HAMTIDE11a. *Upper panels* are amplitudes  $(m^2s^{-1})$ , and *lower panels* are phase lags (deg). The spatial resolution is 0.5°, and the weight for the continuity equation is 10,000



**Fig. 2** Same as Fig. 1 but for  $M_2$ 

ences occur for the diurnal tide K<sub>1</sub>. Moreover, pronounced open-ocean peaks arise on a regular grid in Fig. 1 and are particularly evident in the South Pacific and the Atlantic. Phase values in the original HAM data display similar spikes, corroborating Egbert and Bennett (1996)'s observation of sharp unphysical height changes near observational data for which diagonal covariances are assumed in the assimilation procedure. The gradient operator in the momentum equations then acts to magnify the effect in the flow field. We analyzed longitudinal elevation gradients of K<sub>1</sub> from EOT, finding that jittering features in the tidal heights are specifically pronounced in the aforementioned areas. Such small-scale variations might account for the convergence problems of diurnal tides noted in the previous section; yet, they do not necessarily deteriorate our OTAM results, as these estimates are derived from an average-like global integration.

To assess the accuracy of the deduced volume transports and to select proper continuity equation weights, a comparison of EOT-based OTAM values to estimates from the following assimilation models was performed: HAM, FES2012 (Carrère et al. 2012, hereafter FES), and TPXO7.2 (Egbert and Erofeeva 2002 updated version, hereafter TPX). Amplitude-normalized RMS differences for the harmonic variation of each constituent were computed with respect to TPX, ensuring that the tunable parameters of our solver are not systematically biased toward HAM. Figure 3 shows RMS statistics for eight major tides and includes three EOT solutions computed with weights c = 400, 1000, and 10,000. Other solutions are omitted for readability reasons.

The comparisons clearly argue against the use of a uniform c value for all tides. Instead, different weights per component and tidal species must be assumed to obtain reasonably accurate OTAM estimates. This result presumably reflects the varying spatial characteristics of the flow field as well as the coupling of equatorial motion terms  $(h_{1,2})$ to middle latitudes as against low latitudes in the computation of zonal OTAM estimates  $h_3$ . In detail,  $h_1$  and  $h_2$ exhibit the best agreement with TPX using c = 10,000 for diurnal tides and c between 100 and 400 for semi-diurnal tides. The polar component  $h_3$ , by contrast, requires smaller weights for diurnal tides. Differences in relative RMS can reach 50% or more when adopting c = 400 instead of c = 10,000, which is unacceptable. By and large, the hydrodynamic assimilation models mutually agree within 20%. To achieve inversion results of equally small deviations from TPX, we implemented varying weights for different tidal



Fig. 3 Relative RMS values (dimensionless) of OTAM motion terms with respect to TPX. Weights for the inversion algorithm for the case of EOT as input are shown in the legend.  $\mathbf{a} - \mathbf{c}$  denote x, y, and z components

species as well as equatorial and axial components; see Table 1 for the numerical values. The inconsistencies introduced by that approach are accepted in the frame of this study.

Q1

O1 P1

K1

N2

M2

S2

K2

An intermediate assessment of the inferred motion terms is presented in tabular and graphical form in the online supporting information. EOT-based relative angular momentum estimates blend in well with results from data-assimilative models (HAM, FES, TPX), even though a moderate underestimation of 10–15% is apparent for the equatorial components of M<sub>2</sub>. Conversely, *y* motion terms of our K<sub>1</sub> and O<sub>1</sub> inversions are in slight excess of the harmonics from other models. The following sections will disclose whether these discrepancies have any bearing on the accuracy of rotational predictions from EOT.

 
 Table 1
 Weights for the continuity equation in the final formulation of the inversion algorithm

Туре	Tide	Weight for $h_1, h_2$	Weight for $h_3$
Lunar	$O_1, Q_1$	10,000	100
Solar	P <sub>1</sub>	10,000	400
Luni-solar	K1	10,000	800
Lunar	$M_2$	200	400
Lunar	$N_2$	100	200
Lunar	2N <sub>2</sub>	200	600
Solar	$S_2$	100	2000
Luni-solar	$K_2$	200	1000

#### 3 Minor tides and completion of the ERP model

The altimetric tide atlas EOT11a includes nine gravitational tides in the (sub-)diurnal frequency range ( $Q_1$ ,  $O_1$ ,  $P_1$ ,  $K_1$ ,  $2N_2$ ,  $N_2$ ,  $N_2$ ,  $N_2$ ,  $S_2$ ,  $K_2$ ), but a larger number of spectral components is required for modern-day ERP models (Gipson 1996; Steigenberger et al. 2006; Artz et al. 2011). One possibility to interpolate these minor tides is the response method as described in Munk and Cartwright (1966). The underlying idea is that the admittance, i.e., the ratio between the tidal height and the associated TGP, is a smooth function of frequency. This assumption is also used by the orthotide formulation (Groves and Reynolds 1975) and validated from an OTAM perspective below.

#### 3.1 Admittance relationships

Given the linearity between tidal elevation and tidal angular momentum, admittance interpolation may be also directly applied on OTAM values. Using global momentum integrals from FES tidal heights and currents, we have assessed the smoothness of this OTAM admittance as a function of frequency. Ratios between x, y, and z OTAM versus TGP are plotted against the tidal period for 12 semi-diurnal and five diurnal constituents in Fig. 4. The TGP was again taken from the HW95 catalog. Evidently, diurnal tides display a flat admittance variation with frequency for both mass and motion terms, but the relationship is much more erratic in the semi-diurnal band. Discontinuities in terms of OTAM partic-



Fig. 4 OTAM interpolation using admittance. Shown are ratio values of OTAM (*upper row*: mass term, *lower row*: motion term) and tidal potential for semi-diurnal (*left*) and diurnal (*right*) tides. Angu-

lar momentum values are from FES, and tidal potential amplitudes are from the HW95 catalog. Units:  $10^{25}$  kg m<sup>2</sup>/s

ularly arise for the motion component of the  $2N_2$  tide. The admittance function of the mass term is even rougher, indicating that the assumption of smoothness for semi-diurnal tides is questionable at least for degree-two spherical harmonics of the elevation field. Yet, in the absence of a physically justified way to overcome this complication, we employed a standard linear interpolation of OTAM values to the minor tidal frequencies of interest. The pivot waves used for the interpolation were taken—if available—in accordance with the IERS Conventions (Table 6.7 in Petit and Luzum 2010); see Madzak (2015) for further details.

A total set of 19 minor tides—down to 1% of the TGP of  $M_2$ —were inferred through admittances to complement the nine major constituents of EOT11a. Selected numerical results for these minor tides are provided in Table 2. In detail, OTAM mass estimates were calculated and compared using two approaches: (1) classically, i.e., through interpolating tidal heights of the ocean tide model EOT and subsequent integration; and (2) alternatively, through a direct interpolation of x, y, and z OTAM between the EOT major tides. The RMS values in Table 2 indicate that the difference between the two approaches is well below 1% for the majority of minor tidal constituents. In individual cases, the interpolation results can be compared with hydrodynamically modeled solutions from the FES atlas (last column of Table 2). These RMS differences are in the order of 20% of the OTAM signals, suggesting that input data for minor tides assume a more critical role than the eventual treatment of admittances.

Polar motion and  $\Delta UT1/\Delta LOD$  coefficients for the 28 tidal constituents in our final model were computed using the transfer functions given below and are presented in Tables 5 and 6 in the appendix.

## 3.2 Excitation formalism

Let  $\Delta \hat{I}(\omega)$  be the inertia tensor increment of a particular tide in the frequency domain, and let  $\Omega \Delta \hat{I}(\omega)$  be the mass term of OTAM, scaled to the same units as the motion term or relative angular momentum  $\hat{h}(\omega)$ . In the presence of these perturbations and under the conservation of angular momentum in the solid Earth–ocean system, the linearized Liouville equations for a two-layer Earth dictate the pole position changes  $\hat{p}(\omega)$ (Sasao and Wahr 1981; Gross 1993)

$$\hat{p}(\omega) = \left(2.554 \times 10^{-4} \frac{\Omega}{\omega_{\text{FCN}} - \omega} + 2.686 \times 10^{-3} \frac{\Omega}{\omega_{\text{CW}} - \omega}\right) \frac{\Omega \Delta \hat{I}(\omega)}{A \Omega \tau} + \left(6.170 \times 10^{-4} \frac{\Omega}{\omega_{\text{FCN}} - \omega} + 1.124 \frac{\Omega}{\omega_{\text{CW}} - \omega}\right) \frac{\hat{h}(\omega)}{A \Omega},$$
(11)

**Table 2** Comparison of mass terms from (1) admittance interpolationof tidal elevations and (2) admittance interpolation of OTAM for 19minor constituents

Tide	OTAM <sup>a</sup>	RMS (rel.)	W.r.t. FES
165.565	9.0	7.3 (0.2)	
145.545	11.7	9.3 (0.1)	
275.565	1.8	7.6 (0.5)	
255.545	1.8	7.6 (0.4)	
247.455 (v <sub>2</sub> )	1.3	9.1 (0.7)	245.6 (22.1)
175.455 (J <sub>1</sub> )	3.7	3.8 (0.2)	563.8 (20.8)
155.655 (M <sub>1</sub> )	4.1	2.6 (0.1)	
237.555 (µ <sub>2</sub> )	1.4	22.3 (3.8)	660.7 (70.1)
265.455 (L <sub>2</sub> )	1.3	5.7 (0.5)	188.2 (14.4)
272.556 (T <sub>2</sub> )	1.3	5.4 (0.5)	463.4 (37.2)
185.555 (OO <sub>1</sub> )	2.0	2.7 (0.2)	
137.455 ( <i>ρ</i> <sub>1</sub> )	2.7	3.4 (0.2)	
135.645	2.8	4.0 (0.2)	
127.555 (σ <sub>1</sub> )	2.8	6.8 (0.3)	
155.455	1.5	0.9 (0.1)	
165.545	1.3	1.1 (0.2)	
185.565	1.3	1.7 (0.2)	
162.556 (π1)	1.3	1.0 (0.2)	
125.755 (2Q1)	2.4	6.1 (0.3)	

Values denote RMS differences of the two approaches, computed as mean over *x*, *y*, and *z* for each tidal line. RMS differences between EOTand FES-based OTAM estimates are also specified (fourth column). Units: OTAM in column two in  $10^{23}$  kg m<sup>2</sup>/s, RMS in  $10^{20}$  kg m<sup>2</sup>/s, relative RMS in percent

<sup>a</sup> Computed from interpolated tidal elevations

where  $\tau = \Omega^2 R_E^5 / (3GA)$ , *G* is the universal gravitational constant,  $\omega_{\text{FCN}}$  and  $\omega_{\text{CW}}$  are the frequencies of the Free Core Nutation (FCN) and the Chandler wobble, respectively, and *A* designates the principal equatorial moment of inertia. Equation (11) allows for a straightforward extraction of the prograde and retrograde polar motion components for any tidal constituent.

Changes in length-of-day (in seconds per nominal solar day,  $\text{LOD}_0 = 86400$  s) relate to the axial mass and motion terms of OTAM  $\Omega \Delta I_{33}$  and  $h_3$  through (Gross 1993)

$$\Delta \hat{\text{LOD}}(\omega) = \frac{1}{C_m \Omega} \left( \hat{h}_3(\omega) + 0.756 \ \Omega \ \Delta \hat{I}_{33}(\omega) \right), \quad (12)$$

where  $C_m$  is the principal axial moment of inertia of the Earth's crust and mantle. To convert Fourier coefficients of  $\Delta$ LOD to those of  $\Delta$ UT1, the following equation can be used:

$$\Delta \hat{\text{UT1}}(\omega) = -i \left(\omega \text{LOD}_0\right)^{-1} \Delta \hat{\text{LOD}}(\omega).$$
(13)

## 4 Validation using VLBI observations

In this section, we aim to test the quality of our tide-predicted ERP variations within the analysis of nearly 20 years of VLBI observations. We readily admit that such a single-technique validation is not fully rigorous and should be ideally complemented by a reprocessing of long time-span GPS data, which exhibit an arguably greater sensitivity to sub-diurnal terrestrial rotations than VLBI. However, given the ocean modeling character of our study as well as the computational demands of a full GPS analysis, we leave the latter as a potential field of activity for seasoned GNSS analysts. Note also that any GPS-based assessment of tidal ERP signals needs to contend with orbit resonances and associated systematic errors, e.g., for diurnal and near-diurnal polar motion variations (S. Desai, personal communication, 2016).

Along with the short period ERP model derived in the previous section, three other models were used in separate (but identically configured) VLBI solutions. These datasets are (1) FES-based ocean tidal ERP corrections (called VFES); (2) the model recommended by the IERS Conventions 2010 (hereafter IERS); and (3) an empirical standard deduced from GPS and VLBI observations (Artz et al. 2012), subsequently called VIGG. The one based on EOT is denoted by VEOT. Figure 5 shows a priori ERP harmonics for all four datasets and demonstrates that their level of mutual agreement varies depending on the partial tide, as, e.g., P<sub>1</sub> and Q<sub>1</sub> harmonics agree well, but S2 and M2 show larger differences throughout. Interestingly, the two models based on current ocean tide knowledge, VFES and VEOT, are occasionally more inconsistent with each other than with any of the remaining datasets; cf., e.g., results for N2 and S2 where VEOT closely matches the VIGG values excepting the retrograde band. This and similar findings based on TPX and HAM in Böhm (2012) suggest that even though recent ocean tide models have achieved a good pointwise agreement in terms of elevation, their skill in reflecting the broad-scale (degree-two) features of the tide and the flow field in particular is still far from being satisfactory. The larger uncertainties attached to the motion term contribution are also evident from the spread of OTAM estimates in the online supporting information.

## 4.1 Input data and processing

Available geodetic VLBI observing sessions were subsetted to a total of 1162 24-h sessions from January 1994 through October 2013 according to some rough measures of the network suitability for the estimation of ERP. These input data include R1 and R4 experiments, i.e., weekly VLBI experiments carried out on Mondays and Thursdays, as well as the continuous VLBI campaigns CONT02, CONT05, CONT08, and CONT11. Hourly values of polar motion and  $\Delta$ UT1 were obtained in a two-step solution strategy, in which daily



Fig. 5 Phasor diagrams of ERP values for eight major tidal constituents from four input datasets (VFES, IERS, VIGG, VEOT). Subscripts c and s for polar motion p and  $\Delta$ UT1 denote cosine and sine amplitudes, while

'+' and '--' discriminate prograde and retrograde components of p; see Petit and Luzum (2010) for the respective phase conventions. Note also the different scalings of x and y axes in each panel

EOP series were estimated in a first run to derive accurate nutation offsets for all subsequent analyses. Such a standard time series approach allows for an a posteriori fitting of tidal constituents, but neglects possible stochastic relations between rotations and other parameters of the VLBI analysis. A more rigorous method of estimating tidal ERP variations is to express polar motion and  $\Delta$ UT1 perturbations as sum of sinusoids in the observation equations and solve for the harmonic coefficients simultaneously with other parameters in one global inversion (Gipson 1996). Notwithstanding these conceptual differences, tidal ERP estimates from both approaches have been shown to agree well within the threefold sigma level of any constituent (Böhm 2012).

All processing steps were carried out using the Vienna VLBI Software (Böhm et al. 2012) based on guidelines of the International VLBI Service for Geodesy and Astrometry

Parameter	A priori model	Estimated	Constraint
Ephemerides	JPL421	_	_
CRF	ICRF2	_	_
TRF	VieTRF13 (Krásná et al. 2014)	Offset <sup>a</sup>	NNT/NNR
Polar motion	Pre-solution <sup>b</sup> [IERS 08 C04] + respective high-frequency models (ocean tides, libration)	Hourly [offset <sup>a</sup> ]	1 mas after 1 h
$\Delta UT1$	Pre-solution <sup>b</sup> [IERS 08 C04] + respective high-frequency models (ocean tides, libration)	Hourly [offset <sup>a</sup> ]	1 ms after 1 h
PrecNut.	Pre-solution <sup>b</sup> [IAU2006/2000A]	- [Offset <sup>a</sup> ]	_
Hydrost. delay	Saastamoinen, VMF1	_	_
Wet delay <sup>c</sup>	_	Hourly	1.5 cm after 1 h
Trop. gradients	_	6-Hourly	0.5 mm after 6 h
Clocks <sup>d</sup>	_	Hourly	1.3 cm after 1 h

Table 3 VLBI analysis configuration for the ERP model validation

<sup>a</sup>Per session

<sup>b</sup>In a first step, EOP were determined without estimating high-frequency ERP. Values in square brackets show the a priori source model in this pre-solution

<sup>c</sup>The wet delay was estimated as wet zenith delay

<sup>d</sup>Clocks were estimated as offset + linear trend + quadratic term per clock with respect to a reference clock

Table 4Comparison ofVLBI-derived residuals for thefour ERP models used in thepresent study

	$x_p$ (µas)			$y_p$ (µas)			$\Delta UT1 (\mu s)$			
Model	$\overline{\text{RMS}(d)}$	RMS (s)	$\overline{\Delta \varepsilon}$	RMS(d)	RMS (s)	$\overline{\Delta \varepsilon}$	RMS(d)	RMS (s)	$\overline{\Delta \varepsilon}$	
VFES	8.5	7.7	-0.52	9.0	9.2	-1.30	0.74	0.80	0.01	
IERS	6.9	6.5	1.37	7.8	10.0	-0.70	0.80	0.93	-0.21	
VIGG	7.1	7.6	_	6.9	8.6	_	0.60	0.82	-	
VEOT	8.3	9.5	0.64	7.2	9.1	-0.07	0.81	0.70	0.06	

RMS values for diurnal (d) and semi-diurnal (s) bands were computed from  $\sqrt{A(\varepsilon)^2}/2$ , where A are Fourier coefficient amplitudes of ERP residuals  $\varepsilon$  and the overbar denotes the arithmetic mean.  $\Delta \varepsilon$  are median values of Fourier spectrum differences with respect to VIGG

(IVS). Details of our analysis configuration and the treatment of various parameters can be found in Table 3. Note in particular that the IERS prior for spin and polar motion libration in our analysis is sufficiently accurate ( $\sim$ 3 µas for prograde polar motion) to be excluded from the discussion of residuals below; cf. Brzeziński (2003) and references therein. For all three ERP components, the spectral estimates at 47 tidal frequencies were derived within a least-squares adjustment of discrete hourly time series. From the 71 (sub-)diurnal frequencies in the IERS Conventions 2010 (Tables 8.2a and 8.2b), we removed 24 tidal lines that are separated in frequency from other constituents by 1/18 cycle per year or less.

## 4.2 Discussion of tidal ERP residuals

Fitted tidal residuals  $\varepsilon$  relative to the background model are small for any suitable high-frequency prior and serve as measure by which the performance of our four ERP models can be compared. Specifically, we have mapped the amplitudes A ( $\varepsilon$ ) for 24 diurnal and 23 semi-diurnal tidal residuals to RMS-like composite statistics and specify the numerical results in Table 4. Being optimized toward space geodetic ERP determinations, VIGG naturally produces lowest RMS scores in an average sense, but the superiority of that model is not as clear-cut as one might suspect. To some extent, the limited sensitivity of VLBI to high-frequency ERP variations conceals the goodness of the GPS-dictated VIGG solution; yet, we note that  $x_p$  RMS values from the IERS standard (6.5–6.9 µas) are somewhat below the VIGG differences (7.1–7.6  $\mu$ as). For  $\Delta$ UT1 oscillations, diurnal signals appear to be appropriately accounted for by VIGG to a level of 0.60 µs, but ocean tide-predicted ERP models perform exceedingly well for semi-diurnal variations in  $\Delta UT1$ . The RMS reduction from 0.93  $\mu$ s for the IERS run to 0.70  $\mu$ s for VEOT is particularly encouraging and in fact above the 95% confidence level in a standard t test for 23 semi-diurnal samples.

We now complement and extend the numerical values in Table 4 by analyzing the adjustment results in graphical



Fig. 6 Differences of amplitude spectra of ERP residuals from VLBI analysis. Smaller values indicate larger ERP residuals compared to the reference (VIGG). *First row:*  $x_p$ , *second row:*  $y_p$ , *third row:*  $\Delta$ UT1. VEOT markers filled with a *white circle* indicate that the respective tide is not included in Tables 5 and 6

form. In detail, amplitudes  $A(\varepsilon)$  from three of the four ERP solutions were subtracted from the fourth—the reference solution. The larger the residual coefficients of a model, the more negative is the deviation  $\Delta \varepsilon$  with respect to the reference. Figure 6 illustrates these differential spectra for both pole coordinates and  $\Delta UT1$  in the diurnal and semi-diurnal frequency band. VIGG is taken as reference, so that values below zero imply smaller residuals in the VIGG run. In agreement with Table 4 and the median values of  $\Delta \varepsilon$  specified therein, a tendency toward negative spectral differences is indeed observed for  $y_p$  variations and the diurnal  $\Delta$ UT1 component.

VEOT, the model derived from empirical ocean tides, performs well in most aspects, but produces relatively large polar motion residuals of about  $-15 \,\mu$ as in the O<sub>1</sub> band as well as a pronounced negative  $M_2$  peak of  $-35 \mu$  as in the direction of  $x_n$ . The latter finding is consistent with the moderate underestimation of OTAM motion terms noted in Sect. 2.5. A similar deficiency might have been expected for  $M_2$  in the  $y_n$ spectrum, though this particular term appears to be governed by an FCN-amplified mass term that is accurately represented by EOT11a, with an amplitude much smaller than in the IERS source model  $(3.5 \times 10^{24} \text{ kg m}^2/\text{s} \text{ as against } 4.3 \times 10^{24} \text{ kg m}^2/\text{s}$ in Chao et al. 1996). Residuals from the VFES run approach small values for semi-diurnal pole variations, but diurnal tide corrections at O<sub>1</sub> and K<sub>1</sub> in this model deviate from VIGG in the order of 10-20 µas. That polar motion oscillations at K1 are poorly accounted for by the FES-based tide conforms with the high-latitude regional mass anomalies noted by Stammer et al. (2014) in a comparison of FES to intersatellite ranging data.

Sizable shortcomings of the IERS model are most obvious in the zonal direction, both for diurnal and semi-diurnal  $\Delta$ UT1 variations and minor tidal contributions in particular, e.g., at  $\nu_1$ , Oo<sub>1</sub>, So<sub>1</sub>, J<sub>1</sub>, L<sub>2</sub>, or  $\mu_2$ . For some of these terms, the FES atlas—which models minor tides explicitly—produces a close match with the space geodetic VIGG reference. This result suggests that the handling of minor tidal harmonics in the IERS standard is not without flaws and that modern ocean tide models with a larger number of small constituents can redress such imperfections. Note also that the assumption of zero amplitudes for unconsidered minor tides in the 28-line VEOT model tends to lead to a reduction of residuals relative to VIGG (white dots in Fig. 6).

For our VLBI analysis, VEOT appears to be the most suitable prior model in the semi-diurnal  $\Delta$ UT1 component, whereas its skill in the diurnal band is affected by errors of the O<sub>1</sub> and K<sub>1</sub> tides. By virtue of the applied admittance interpolation (Sect. 3), these deficiencies are also transferred to many of the surrounding minor tides. At the O1 line, which has been a notorious anomaly in preceding ERP studies, the relative VEOT residual is  $-2.3 \,\mu s$  and in considerable excess of  $\Delta \varepsilon$ from VFES  $(-1.0 \,\mu s)$ . While the predominant excitation of O<sub>1</sub> tidal variations in Earth's spin comes from the motion component of OTAM, we recall that non-conservation of mass in EOT (and also in FES) is particularly pronounced for O<sub>1</sub>; cf. Sect. 2.3. Evidently, the incorporation of global factors such as mass conservation into ocean tide models along with an improved account of the volume transports in low latitudes are pivotal steps toward a credible standard of tide-induced diurnal  $\Delta UT1$  variations.

# 5 Summary and conclusions

The present work has explored one method by which the longstanding necessity for an updated model of sub-diurnal

Earth rotation variations due to dynamic ocean tides might be met. Elevation fields of nine major constituents were fixed to a recent empirical ocean tide model, and barotropic volume transports were estimated based on a least-squares inversion of linearized shallow water equations in the fashion of Ray (2001). The recent progress of altimetry-constrained tide models in resolving open-ocean tidal heights to the accuracy of 1 cm (Stammer et al. 2014) prompted high initial expectations for the quality of ERP predictions from this inversion scheme; yet, our study has led to the conclusion that angular momentum changes in the ocean-derived either from the present approach or via data-assimilative forward modelsstill include uncertainties that make the development of such an ERP model to the 5-10 µas accuracy level difficult. In this respect, the ambitious goal of reducing the ERP model errors by approximately one order of magnitude (cf. Griffiths and Ray 2013) could not be met.

Concerning the modeling approach adopted here, errors of motion terms and hence the numerically deduced tidal currents play a critical role and require knowledge of dissipative terms and the discrete model bathymetry. Comparisons in Madzak (2015) have shown that a simple replacement of the regridded 0.5° GEBCO chart with Shuttle Radar Topography Mission (SRTM) data in an otherwise unchanged least-squares solution alters motion terms in the order of 10%. Regarding dissipational processes, wave scattering into internal tides near major topographic features is now routinely included in barotropic tide models in parametrized form (Green and Nycander 2013), but no consensus exists about which conversion scheme performs best over a large number of constituents. Employing such drag terms in our linearized momentum equations is straightforward, yet when doing so, one is faced with the task of delimiting them from the bottom friction stresses of the very same linear form.

In addition to unavoidable dynamical approximations, the inversion algorithm used in this study also imposes numerical challenges. Input elevation charts may contain artificial gridding features that affect the convergence of the iterative least-squares solver, while too small frictional parameters in the open ocean generally result in a poorly conditioned coefficient matrix. Above all, the estimated volume transports critically depend on the the relative weight of the continuity constraint against the momentum equations, and credible motion terms were in fact only obtained after having tuned that value per constituent. A dynamically more consistent weighting approach is targeted for future implementations, possibly along with tests of other sparse matrix solvers than the built-in MATLAB algorithm.

By and large, and from the viewpoint of our VLBI validation, the deduced high-frequency ERP model has led to improvements over the current IERS standard for individual major tides (K<sub>1</sub> in  $y_p$ , P<sub>1</sub>, Q<sub>1</sub>, K<sub>2</sub>) and several adjacent minor tides accounted for through admittance interpolation. In contrast, our results for O<sub>1</sub> and the principal M<sub>2</sub> tide in polar motion are admittedly unsatisfying and require more detailed examination. Full hydrodynamic tide solutions from dataconstrained assimilation models appear to be a viable tool to mitigate these problems, and Fig. 6 suggests that advanced fine-resolution tidal atlases such as FES do indeed show signs of convergence with VLBI-derived ERP variations in the semi-diurnal band. Modeling the rotational influences of diurnal tides is still a delicate matter, though, especially for polar motion oscillations that couple to higher latitudes in which altimetric observations are increasingly sparse, the bottom topography is known with limited accuracy, and regional cancellation effects (e.g., due to the Antarctic Kelvin wave) are very pronounced. We are, however, confident that forthcoming altimeter missions and dynamically consistent tide modeling will eventually allow for a precise a priori description of high-frequency Earth rotation variations.

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# Appendix: Coefficients of the ERP model

Tables 5 and 6 present the diurnal and semi-diurnal polar motion and UT1/LOD variations derived in this study.

 Table 5
 Ocean tidal variations in polar motion as deduced from EOT. In addition to nine major tides initially included in the altimetric model, OTAM values of 19 minor tides were inferred from linear admittance interpolation

Tide	Del	aunay ar	guments				Doodson	Period (h)	$x_p \sin(\mu as)$	$x_p \cos(\mu as)$	$y_p \sin(\mu as)$	$y_p \cos(\mu as)$
	γ	l	l'	F	D	Ω	- number					
SQ <sub>1</sub>	1	-2	0	-2	0	-2	125.755	28.0062	0.7	4.8	-4.8	0.7
$\sigma_1$	1	0	0	-2	-2	-2	127.555	27.8484	0.9	5.7	-5.7	0.9
	1	-1	0	-2	0	-1	135.645	26.8728	1.9	5.8	-5.8	1.9
Q1	1	-1	0	-2	0	-2	135.655	26.8684	10.3	30.5	-30.5	10.3
$\rho_1$	1	1	0	-2	-2	-2	137.455	26.7231	2.0	5.6	-5.6	2.0
	1	0	0	-2	0	-1	145.545	25.8234	12.2	27.4	-27.4	12.2
O1	1	0	0	-2	0	-2	145.555	25.8193	65.1	146.1	-146.1	65.1
	1	1	0	-2	0	-2	155.455	24.8492	-1.4	-3.8	3.8	-1.4
$M_1$	1	-1	0	0	0	0	155.655	24.8332	-4.0	-10.5	10.5	-4.0
$\pi_1$	1	0	-1	-2	2	-2	162.556	24.1321	1.9	3.3	-3.3	1.9
<b>P</b> <sub>1</sub>	1	0	0	-2	2	-2	163.555	24.0659	28.1	40.3	-40.3	28.1
	1	0	0	0	0	1	165.545	23.9380	2.2	3.1	-3.1	2.2
K1	1	0	0	0	0	0	165.555	23.9345	-113.0	-158.0	158.0	-113.0
	1	0	0	0	0	-1	165.565	23.9310	-15.3	-21.4	21.4	-15.3
$J_1$	1	1	0	0	0	0	175.455	23.0985	-11.9	-6.4	6.4	-11.9
001	1	0	0	2	0	2	185.555	22.3061	-11.2	-1.9	1.9	-11.2
	1	0	0	2	0	1	185.565	22.3030	-7.2	-1.2	1.2	-7.2
$2N_2$	2	-2	0	-2	0	-2	235.755	12.9054	-6.6	-0.9	1.1	4.0
$\mu_2$	2	0	0	-2	-2	-2	237.555	12.8718	-8.0	-1.1	1.2	4.8
$N_2$	2	-1	0	-2	0	-2	245.655	12.6583	-49.4	-10.3	4.3	32.5
$\nu_2$	2	1	0	-2	-2	-2	247.455	12.6260	-9.6	-1.8	1.0	6.1
	2	0	0	-2	0	-1	255.545	12.4215	10.6	0.9	-1.6	-6.5
$M_2$	2	0	0	-2	0	-2	255.555	12.4206	-282.1	-24.6	43.6	174.0
L <sub>2</sub>	2	1	0	-2	0	-2	265.455	12.1916	8.4	-1.1	-2.3	-5.3
$T_2$	2	0	-1	-2	2	-2	272.556	12.0164	-8.2	2.4	3.1	4.7
$S_2$	2	0	0	-2	2	-2	273.555	12.0000	-132.3	53.1	56.1	84.3
$K_2$	2	0	0	0	0	0	275.555	11.9672	-38.0	12.7	15.5	21.4
	2	0	0	0	0	-1	275.565	11.9664	-11.4	3.8	4.6	6.3

 Table 6
 Ocean tidal variations in UT1 and LOD as deduced from EOT. In addition to nine major tides initially included in the altimetric model, OTAM values of 19 minor tides were inferred from linear admittance interpolation

Tide	Dela	unay argu	iments				Doodson	Period (h)	UT1	UT1		LOD	
	γ	l	l'	F	D	Ω	number		sin (µs)	cos (µs)	sin (µs)	cos (µs)	
SQ <sub>1</sub>	1	-2	0	-2	0	-2	125.755	28.0062	0.9	-0.3	-1.7	-4.6	
$\sigma_1$	1	0	0	-2	-2	-2	127.555	27.8484	1.0	-0.4	-2.1	-5.4	
	1	-1	0	-2	0	-1	135.645	26.8728	0.8	-0.5	-2.7	-4.7	
Q1	1	-1	0	-2	0	-2	135.655	26.8684	4.4	-2.5	-14.2	-24.6	
$\rho_1$	1	1	0	-2	-2	-2	137.455	26.7231	0.8	-0.5	-2.7	-4.4	
	1	0	0	-2	0	-1	145.545	25.8234	2.9	-2.5	-14.5	-16.8	
O1	1	0	0	-2	0	-2	145.555	25.8193	15.3	-13.3	-77.6	-89.5	
	1	1	0	-2	0	-2	155.455	24.8492	-0.4	0.3	1.8	2.2	
$M_1$	1	-1	0	0	0	0	155.655	24.8332	-1.0	0.9	5.3	6.3	
$\pi_1$	1	0	-1	-2	2	-2	162.556	24.1321	0.3	-0.2	-1.4	-1.9	
P1	1	0	0	-2	2	-2	163.555	24.0659	4.6	-2.9	-18.2	-29.0	
	1	0	0	0	0	1	165.545	23.9380	0.3	-0.2	-1.3	-1.9	
K <sub>1</sub>	1	0	0	0	0	0	165.555	23.9345	-15.3	9.7	61.4	96.5	
	1	0	0	0	0	-1	165.565	23.9310	-2.1	1.3	8.3	13.1	
$J_1$	1	1	0	0	0	0	175.455	23.0985	-0.7	0.1	0.6	4.3	
$OO_1$	1	0	0	2	0	2	185.555	22.3061	-0.2	-0.3	-1.7	1.7	
	1	0	0	2	0	1	185.565	22.3030	-0.2	-0.2	-1.1	1.1	
2N <sub>2</sub>	2	-2	0	-2	0	-2	235.755	12.9054	-0.4	-0.4	-5.0	5.2	
$\mu_2$	2	0	0	-2	-2	-2	237.555	12.8718	-0.5	-0.5	-5.9	6.3	
$N_2$	2	-1	0	-2	0	-2	245.655	12.6583	-2.7	-2.1	-24.9	32.1	
$\nu_2$	2	1	0	-2	-2	-2	247.455	12.6260	-0.5	-0.4	-4.7	6.4	
	2	0	0	-2	0	-1	255.545	12.4215	0.6	0.3	3.5	-7.4	
$M_2$	2	0	0	-2	0	-2	255.555	12.4206	-16.5	-8.0	-96.7	200.3	
L <sub>2</sub>	2	1	0	-2	0	-2	265.455	12.1916	0.4	0.3	3.1	-4.5	
$T_2$	2	0	-1	-2	2	-2	272.556	12.0164	-0.4	-0.2	-2.1	5.4	
$S_2$	2	0	0	-2	2	-2	273.555	12.0000	-8.7	-1.8	-22.2	109.9	
K <sub>2</sub>	2	0	0	0	0	0	275.555	11.9672	-2.2	-0.6	-7.2	27.7	
	2	0	0	0	0	-1	275.565	11.9664	-0.7	-0.2	-2.2	8.3	

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